

MININ, I.N.

Radiation pressure and the dynamics of planetary nebulae [with summary
in French]. Vop.kosm. 6:211-220 '58. (MIRA 11:10)
(Cosmic physics) (Nebular hypothesis)

AUTHORS: Shifrin, K. S. and Minin, I. N. SOV/49-59-1-15/23
 TITLE: Non-Horizontal Visibility Below a Continuous Layer
 of Cloud (Negorizental'naya vidimost' pri sploshnoy
 oblachnosti)
 PERIODICAL: Izvestiya Akademii Nauk SSSR, Seriya Geofizicheskaya,
 1959, Nr 1, pp 131-138 (USSR)
 ABSTRACT: Visibility in the atmosphere below a continuous layer
 of cloud is considered. A formula for calculating
 the contrast K is given, p 131, where h - height of
 the observation point, Θ - angle of observation,
 $B_o(\lambda)$ - brightness of an object on the Earth's surface,
 the brightness of which is $B_o(\lambda)$, λ - wavelength,
 $D(\lambda)$ - brightness of the haze, $\tau_o^h(\lambda)$ - optical thickness
 of the air layer, F_1, F_2, F_3 - energy streams from the
 object, Earth's surface and haze respectively,
 L - distance from the object, ϵ - limit of sensitivity
 of the visibility meter. The non-horizontal distance
 L can be found from the expression

$$K(L) = \epsilon$$

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Non-Horizontal Visibility Below a Continuous Layer of Cloud

or

$$L = K^{-1}(\epsilon).$$

The coefficient of diffusion is given by Eq.(1), the thickness of the air layer by Eq.(2) and the optical thickness by Eq.(3). The brightness of the air haze can be calculated from Eq.(4) (Ref 3) or Eq.(5) where $J(\tau'_0, \Theta) = \sigma(\Theta)$. In the general case, this equation can be written in the form of Eq.(6). The conditions satisfying Eq.(6) are shown in Fig.1 and Table 1. The brightness of the cloud haze can be calculated from Eq.(7) where D is found experimentally (Ref 1). A mean D can be calculated from Eq.(8). Thus $J(\Theta)$ becomes simplified as is shown in Eq.(9). The integral of this equation can be evaluated and presented as Eq.(10) when correction for height of the Sun $i = 20$ to 80° is applied. The spectral brightness of the Earth's surface can be calculated from Eq.(11) where $I_0(\lambda)$ - stream of parallel rays from the cloud. That part of the light which falls from a portion of cloud at an angle of $d\omega$ can be found from Eq.(12).

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Non-Horizontal Visibility Below a Continuous Layer of Cloud

Substituting $J(\tau'_0, \Theta)$ from Eq.(4), the Eq.(13) is obtained, while the total illumination by the whole sky can be found from Eqs.(14), (15) and (16) where $\Phi(\lambda)$ is spectral illumination. The tables of Φ were made for the values of τ ranging from 0.00 to 0.50 and for A from 0.0 to 1.0 (Table 2). As can be seen, the value of Φ can be considered as constant and equal to about 2. The calculation of the brightness of the Earth's surface can also be based on its own spectral illumination $E(\lambda)$. The ratio of $E(\lambda)/J(\lambda)$ can be found from Eq.(18), thus the brightness of the cloud can be expressed as Eq.(19) and the coefficient ϵ calculated from Eqs.(20)-(22). How the value of ϵ depends on λ can be shown in an example for $S_0 = 20$ km, $\tau_0 = 0.3$, $\Theta = 60^\circ$. Taking $A = 0.2$, the following computation can be performed:

$$(1 - A) \left(1 + \frac{3}{2} \cos \Theta \right) + 2A = 1.8,$$

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$$\phi(\tau^0, A) = 0.9 [4 + (3 - x_1) 0.7 \tau_0] = 3.6 + 1.3 \tau_0.$$

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Non-Horizontal Visibility Below a Continuous Layer of Cloud

The result is shown in Table 3. It can be seen that the deformation of the spectral curve of brightness distribution, downwards from the cloud base of 1.53 km, does not exceed 12%. When the cloud base is relatively high (i.e. 2-4 km in the summer) a correction should be applied in Eq.(6). This can be found from Eqs.(23) and (24) where η and τ are found experimentally for the values of θ equal 45 and 60° (Table 4). Similar tables can be made for various η and τ . Thus, knowing τ it will be easy to determine the decrease in illumination of an object and the Earth's surface or of the brightness of the haze, thus determining the contrast K.

There are 1 figure, 4 tables and 6 references, all of which are Soviet.

ASSOCIATION: Glavnaya geofizicheskaya observatoriya im.A.I.Voyeykova
(Main Geophysical Observatory imeni A.I.Voyeykov)

SUBMITTED: September 23, 1957

Card 4/4

21(8)

AUTHOR: Minin, I.N.

SOV/43-59-13-13/16

TITLE: On the Solution of Instationary Problems of Radiation Transfer Theory

PERIODICAL: Vestnik Leningradskogo universiteta, Seriya matematiki, mekhaniki i astronomii, 1959, Nr 13(3), pp 137-141 (USSR)

ABSTRACT: Joining the papers of V.V.Sobolev [Ref 1,2,3] the author investigates instationary problems of the radiation transfer theory. The investigation is carried out with means of Laplace transformations and with statistical methods of V.V.Sobolev. The author obtains the radiation intensity on the boundary as well as in the medium itself. Partly the known solutions of the corresponding stationary problems are used. The author determines especially the probability of the quantum reflection at a semi-infinite medium in the time interval u to $u+du$. There are 4 Soviet references.

SUBMITTED: February 22, 1958

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10(1),10(4)

AUTHOR: Mikheyev, A.S.

307/43-59-13-14/16

TITLE: Equations of Gas Dynamics in the Case of Axial Symmetry

PERIODICAL: Vestnik Leningradskogo universiteta, Seriya matematiki, mekhaniki i astronomii, 1959, Nr 13(3), pp 142-144 (USSR)

ABSTRACT: The author establishes the motion equations of an axialsymmetric flow of an ideal, compressible fluid. It is assumed that the density is known as an arbitrary function of pressure and flow function. If the velocity depends on the pressure only, then we have a laminar motion. If $v = v_1(p) \cdot F(\psi)$, where p is the pressure and ψ is the flow function, then the investigation can be reduced to the consideration of a potential flow with the velocity $v_1(p)$ and the flow function $\psi_1 = \int F(\psi) d\psi$.

The author mentions L.I.Sedov and Yu.V.Rudnev. There are 3 references, 2 of which are Soviet, and 1 German.

SUBMITTED: August 15, 1958

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S/033/60/037/005/022/024
E032/E514

AUTHOR: Minin, I.N.

TITLE: On the Motion of a Variable Mass Envelope

PERIODICAL: Astronomicheskiy zhurnal, 1960, Vol.37, No.5,
pp.939-940

TEXT: In 1958 Mustel' (Ref.1) considered the motion of an envelope in an inter-stellar medium of density ρ_c assuming that the mass of the envelope continuously increases at the expense of the inter-stellar medium and also as a result of the outflow of matter from the star. The formula which he used for the amount of matter flowing out of the star was $4\pi R_o^2 \rho_o v_o$, where R_o is the radius of the layer from which the outflow takes place, ρ_o is its density and v_o the outflow velocity. The problem was formulated in detail in an earlier paper due to Mustel' (Ref.2). He obtained the following equation of motion for an envelope:

$$\frac{d^2 R_a}{dt^2} = \frac{4\pi R_o^2 \rho_o (v_o - v_a)^2 - 4\pi R_a^2 \rho_c v_a^2}{m_o + 4\pi R_o^2 \rho_o (v_o t - R_a + R_{a_o}) + \frac{4}{3}\pi \rho_c (R_a^3 - R_{a_o}^3)} \quad (1)$$

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On the Motion of a Variable Mass Envelope

where R_a is the radius of the envelope at the time t , v_a is the expansion velocity of the envelope, R_{a0} is the radius of the

envelope at $t = 0$ and m_0 is the initial mass of the envelope. In another paper (Ref.3) Mustel' reported numerical solutions of Eq.(1) obtained with the aid of the electronic computer "Strela". The solutions were obtained subject to the conditions $t = 0, R_a = R_{a0}, v = v_0, m = \frac{4}{3}\pi R_{a0}^3 \rho_c$. In 1954 the present author also discussed this problem and obtained an equation of motion for the envelope which was of the form

$$\frac{d}{dt} (mv) = aV (V - v) \quad (2)$$

where

$$m = m_0 + \frac{4}{3}\pi r^3 \rho + a \int_0^t (V - v) dt \quad (3)$$

In these two equations r and v are the radius and the expansion
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On the Motion of a Variable Mass Envelope

velocity at a time t , aV is the amount of matter flowing out of the the star, V is the outflow velocity, ρ is the density of the inter-stellar medium and m and m_0 are the masses of the envelope at a time t and at $t = 0$ respectively. The initial conditions employed were: $t = 0$, $r = 0$, $v = 0$, $m = m_0$ and a solution of the problem was obtained in a closed form. Bearing in mind that $v = dr/dt$, Eq.(2) can be re-written in the form

$$m \frac{d^2 r}{dt^2} = aV (V - v) - v \frac{dm}{dt} \quad (4)$$

and hence using Eq.(3) one finds that

$$\frac{dm}{dt} = 4\pi r^2 \rho v + a(V - v) \quad (5)$$

Since subject to the above initial conditions

$$\int_0^t (V - v) dt = Vt - r \quad (6)$$

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On the Motion of a Variable Mass Envelope

it follows that substituting Eqs. (3), (5) and (6) into Eq. (4) the final result is

$$\frac{d^2 r}{dt^2} = \frac{a(V - v)^2 - 4\pi r^2 \rho v^2}{m_0 + a(Vt - r) + \frac{4}{3}\pi \rho r^3} \quad (7)$$

Allowing for changes in the notation it is seen that Eq. (7) is identical with Eq. (1). Since the present author obtained a closed solution of the problem, it follows that the numerical procedure employed by Mustel' is unnecessary and an algebraic solution can be obtained. There are 4 Soviet references.

SUBMITTED: April 25, 1960

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MININ, I.N.

Plenum of the Committee on the Physics of Stars and Nebulae.
Vop.kosm. 7:373-375 '60. (MIRA 13:11)
(Astrophysics--Congresses) (Nebulae--Congresses)

S/020/60/133/01/20/070
B014/B011

AUTHOR: Minin, I. N.

TITLE: A Point Source of Light²¹ in an Absorbing Medium Between Parallel Planes

PERIODICAL: Doklady Akademii nauk SSSR, 1960, Vol. 133, No. 1, pp. 74-76

TEXT: The present paper offers a generalization of V. V. Sobolev's study of a point light source placed between parallel isotropically reflecting planes. The method suggested by the said author is applied for the purpose. The system of integral equations obtained is solved by introducing relation (3) for the illuminated planes, instead of formula (2), whereby an Abel integral equation is obtained for (2). Proceeding therefrom the author obtains the exact analytical solution (12) of the problem considered. Finally, the calculation according to formula (12) is discussed. There is 1 Soviet reference.

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A Point Source of Light in an Absorbing
Medium Between Parallel Planes

S/020/60/133/01/20/070
B014/B011

PRESENTED: March 14, 1960, by V. A. Ambartsunyan, Academician

SUBMITTED: March 11, 1960

✓B

Card 2/2

Minin, I. N.

8/020/60/133/03/04/013
B019/B056

AUTHOR: Minin, I. N.

TITLE: The Solution of the Integral Equation of the Coastal Refraction of Electromagnetic Waves 12

PERIODICAL: Doklady Akademii nauk SSSR, 1960, Vol. 133, No. 3, pp. 558 - 560

TEXT: Reference is made in the introduction to a solution by V. A. Fok of the integral equation (1), which occurs in the theory of coastal refraction. Next, the solution (3) of the integral equation (2) is given, which was obtained by a new method developed by V. V. Sobolev (Ref. 3) by a generalization of the work of V. A. Ambartsumyan (Ref. 4). By means of Sobolev's method it is possible to show that the solutions of all equations (2) may be represented by means of a function $\Phi(\tau)$. The determination of this function $\Phi(\tau)$ is dealt with in detail. The application of this method to equation (1) is then discussed. For $\Phi(\tau)$, the author obtains the equations (13) and (14) respectively for different ranges of the parameter α . In conclusion, two special cases are

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The Solution of the Integral Equation of the
Coastal Refraction of Electromagnetic Waves

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B019/B056

investigated where the function $g(\tau)$ describing the wave incident on the coast corresponds to the form $g(\tau, x) = e^{-\tau x}$ and the integral (18), respectively. The latter case represents waves which are superpositions of plane waves. There are 5 Soviet references.

PRESENTED: May 3, 1960, by V. A. Ambartsumyan, Academician

✓

SUBMITTED: April 28, 1960

Card 2/2

20768.

S/043/61/000/001/010/010
C111/C222

24.3200

AUTHOR: Minin, I.N.

TITLE: Diffusion of radiation in a semi-infinite medium with a non-isotropic scattering.I.

PERIODICAL: Leningrad. Universitet. Vestnik. Seriya matematiki, mekhaniki i astronomii, no.1, 1961, 133-143.

TEXT: The author considers the diffusion of the radiation in a semi-infinite medium with a non-spherical indicatrix of scattering with the aid of probability theoretical methods.

Let $p(\tau, \eta', \eta, \varphi' - \varphi)d\omega$ be the probability that a photon absorbed in the optical depth τ from a direction forming the angle arc $\cos \eta'$ with the outer normal of the layers having the azimuth φ' leaves the medium through $\tau = 0$ under the angle arc $\cos \eta$ to the normal and with the azimuth φ (in the solid angle $d\omega$). Let $q(\tau, \eta', \eta, \varphi' - \varphi)d\omega$ be the probability that a photon emitted in the optical depth τ in a direction forming the angle arc $\cos \eta'$ with the outer normal and having the azimuth φ' leaves the medium under the angle arc $\cos \eta$ with the azimuth φ . Several connections between p and q are given, e.g.

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Diffusion of radiation...

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$$q(\tau, \eta', \eta, \varphi' - \varphi) = - \int_{\tau}^{\infty} e^{-\frac{\tau - \tau'}{\eta'}} p(\tau', \eta', \eta, \varphi' - \varphi) \frac{d\tau'}{\eta'} \quad (\eta' < 0). \quad (5)$$

By elimination of q from this connections one obtains for p

$$p(\tau, \eta', \eta, \varphi' - \varphi) = \frac{\lambda}{4\pi} \int_0^{2\pi} d\varphi'' \left[\int_0^1 x(\gamma) d\eta'' \int_0^1 p(\tau', \eta'', \eta, \varphi'' - \varphi) e^{-\frac{\tau - \tau'}{\eta''}} d\tau' - \int_0^1 x(\gamma) d\eta'' \int_0^1 p(\tau', \eta'', \eta, \varphi'' - \varphi) e^{-\frac{\tau - \tau'}{\eta''}} d\tau' \right] + \frac{\lambda}{4\pi} x(\eta) e^{-\frac{\tau}{\eta}}, \quad (7)$$

where

$$\cos \gamma = \eta \eta' + \sqrt{(1 - \eta^2)(1 - \eta'^2)} \cos(\varphi' - \varphi). \quad (8)$$

and $x(\gamma)$ is the indicatrix of scattering. Besides p satisfies the relation

$$\frac{\partial p(\tau, \eta', \eta, \varphi' - \varphi)}{\partial \tau} = - \frac{1}{\eta} p(\tau, \eta', \eta, \varphi' - \varphi) + \int_0^{12\pi} p(\tau, \eta', \eta'', \varphi' - \varphi'') \cdot p(\tau, \eta'', \eta, \varphi'' - \varphi) \frac{d\eta''}{\eta} d\varphi''. \quad (12)$$

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where (following from (7) for $\tau = 0$)

$$p(0, \eta', \eta, \varphi' - \varphi) = \frac{\lambda}{4\pi} \left[x(\tau_1) + \frac{\eta}{\pi} \int_0^{2\pi} d\varphi'' \int_0^1 x(\chi) \xi(\eta'', \eta, \varphi'' - \varphi) \frac{d\eta''}{\eta''} d\varphi'' \right] \quad (13)$$

$\xi(\eta', \eta, \varphi' - \varphi)$ is the reflection coefficient of the medium. Then the intensity of radiation emerging from a medium illuminated by parallel rays is calculated. If $\text{arc cos } \zeta$ is the angle formed by the incident parallel rays with the normal of the boundary of the medium, and if φ_0 is the azimuth of the rays, while $I(0, \eta, \zeta, \varphi - \varphi_0)$ is the intensity of the radiation which leaves the medium under the angle $\text{arc cos } \eta$ to the normal and with the azimuth φ , then

$$I(0, \eta, \zeta, \varphi - \varphi_0) = \pi S \cdot \int_0^\infty e^{-\frac{\tau}{\zeta}} p(\tau, -\zeta, \eta, \varphi_0 - \varphi) \frac{d\tau}{\eta}. \quad (16)$$

(πS is the flow of the radiation through a unit surface lying perpendicular to the rays on $\tau = 0$). With the aid of (16) it is stated:

$$\xi(\eta, \zeta, \varphi - \varphi_0) = \xi(\zeta, \eta, \varphi_0 - \varphi). \quad (20)$$

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Diffusion of radiation...

Particularly the author considers the case where the indicatrix of scattering has the form

$$x(\gamma) = \sum_{i=0}^n x_i P_i(\cos \gamma), \quad (21)$$

where the P_i are Legendre polynomials. The obtained results agree with the older results of V.A.Ambartsunyan (Ref.4: ZhETF, 13, no.9,10, 1943). Finally the intensity of the emitted radiation is considered for the case that there exist several radiation sources. The considered individual cases were already investigated by V.V.Sobolev, Ambartsunyan and others.

There are 6 Soviet-bloc and 1 non-Soviet-bloc references. The reference to the English-language publication reads as follows: S.Ueno. J.math. a.mech., 7, no.4, 1958.

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MININ, I.N.

Optical characteristics of diffuse nebulae. Astron.zhur.
38 no.4:641-646 J1-Ag '61. (MIRA 14:8)

1. Leningradskiy gosudarstvennyy universitet im. A.A. Zhdanova.
(Nebulae)

SOBOLEV, V.V.; MININ, I.N.

Isotropic light scattering in an atmosphere with finite optical thickness. Astron.zhur. 38 no.6:1025-1032 N-D '61. (MIRA 14:11)

1. Astronomicheskaya observatoriya Leningradskogo gosudarstvennogo universiteta im. A.A.Zhdanova.
(Light--Scattering)

MININ, I. N.,

S/560/62/000/014/0012/011

AUTHOR: Sobolev, V. V., and I. N. Minin

TITLE: Light scattering in a spherical atmosphere. I.

PERIODICAL: Akademiya nauk SSSR, *Iskusstvennyye sputniki Zemli*, no. 14, 1962, 7-12

TEXT: Light scattering in an atmosphere consisting of spherical layers (e. g., when the sun is low on the horizon or beneath it) is examined. An approximate solution of equations for the intensity of radiation (I) and the total quantity of radiation (B) is proposed on the basis of a method used by V. V. Sobolev to solve the problem of light scattering in a medium consisting of plane-parallel layers. First order scattering is accounted for precisely, while scattering of higher orders is approximated. Here only the first two components are used in the expansion of the scattering indicatrix in Legendre polynomials. The equations obtained are valid for all relationships of the coefficient of absorption (α) to the distance (r) of an arbitrary point in the atmosphere from the center of the planet.

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Light scattering ...

S/560/62/000/014/001a/011

Two special cases are considered: 1) where α is constant in the atmosphere and 2) where α decreases exponentially with height. Case (1) may be presumed to exist when the sky is totally overcast and case (2), when it is clear. The computations could be simplified if it were assumed that the thickness of the atmosphere is considerably less than the radius of the planet, as is actually the case. Light scattering in the Venusian atmosphere is recognized as a special case. Here the atmosphere consists of two layers: a cloudy layer with an approximately constant α and an underlying gaseous layer with varying α .

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MININ, I.N.

Theory of unsteady diffusion of radiation. Vest. LGU 17 no.19:124-
132 '62. (MIRA 15:10)

(Radiation) (Diffusion)

L 10339-63

ENT(1)/FCC(w)/BDS/ES(v)... AFFTC/ESD-3 Pa-4 GW

PHASE I BOOK EXPLOITATION

SOV/6434

Gorbatskiy, V. G., and I. N. Minin

Nestatsionarnyye zvezdy (Unstable Stars) Moscow, Fizmatgiz, 1963. 355 p.
(Series: Problemy teoreticheskoy astrofiziki) 2000 copies printed.

Editorial Board of the Series: V. A. Ambartsumyan, E. R. Mustel', A. B. Severnyy, and V. V. Sobolev; Ed.: G. S. Kulikov; Tech. Ed.: I. Sh. Aksel'rod.

PURPOSE: This book is intended for astronomers and astrophysicists.

COVERAGE: Unstable stars, including novae, supernovae, Wolf-Rayet, and Be-types, are investigated on the basis of their emission characteristics. Shell dynamics during flareup are examined. The instability of the stars is interpreted chiefly on the basis of the structure of the outer layers, since little data is available on the interior of such stars. Some attention is given to the application of gasdynamics and electrodynamic techniques

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Unstable Stars

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to the study of unstable stars. Cepheid variables are not examined. Chs. 1, 3, 7, 9, 10, and para. 19 were written by V. G. Gorbatskiy and Chs. 2, 4, 5, 6, and 7, by I. N. Minin. The authors thank colleagues at the Department of Astrophysics of Leningrad University and the Department of Stellar Physics of the Crimean Astrophysical Observatory. There are 293 references, including 132 Soviet.

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Foreword

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PART I. NOVAE

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- 1. Typical novae
- 2. Qualitative interpretation of a nova outburst
- 3. Characteristics of individual novae

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ACCESSION NR: AF4003731

8/0293/63/001/002/0227/0234

AUTHOR: Minin, I. N.; Sobolev, V. V.

TITLE: Light scattering in a spherical atmosphere.

SOURCE: Kosmichaskiya issledovaniya, v. 1, no. 2, 1963, 227-234

TOPIC TAGS: atmospheric light scattering, spherical atmosphere, planetary atmosphere, atmospheric layer curvature, light scattering, light reflection, outgoing radiation, atmospheric absorption, atmospheric optical thickness, planet reflected light, homogeneous sphere luminescence.

ABSTRACT: The article is a continuation of the authors' previous work on the scattering of light in a planetary atmosphere which accounts for the curvature of atmospheric layers (V. V. Sobolev, I. N. Minin. Sb. "Iskusstvennyye Sputniki Zemli," vy* p. 14. Izd-vo AN SSSR, 1962, str. 7). In the present article, the case of a constant atmospheric absorption coefficient is considered. An analytical solution is obtained for the basic equation determining the mean intensity of the diffused radiation, J , at a point in the atmosphere, subject to boundary conditions. These conditions assume that there exists no diffused radiation incident upon the atmosphere from

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the outside, and they account for the reflection of light from the planet's surface. The expression for the quantity J of a homogeneous sphere is derived for the optical thickness of the atmosphere, which is large in comparison to the planet dimensions. The result is similar to, but simpler than that obtained by R. G. Giovanelli and J. T. Jefferies (Proc. Phys. Soc., 69, No. 11, 1077, 1956). From the knowledge of J , the ratio B of the radiation coefficient to the absorption coefficient can be derived for any point. The intensity of radiation leaving the atmosphere is then expressed as:

$$I = \int_0^{T_0} B e^{-T_1} dT_1 + I^* e^{-T_0},$$

where T_1 is the range along a ray of light between a point in the atmosphere and the observer, T is the total path traveled by the ray in the atmosphere, and I^* is the intensity of radiation due to reflection from the planet's surface. The integral of the equation is written as $I_1 + \Delta I$, where I_1 is the intensity due to first order scattering and ΔI represents higher

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orders. For the case when the atmosphere can be approximated by a homogeneous sphere and the observer is at a far field, the coordinates of any point are easily expressed in terms of T_1 , and an explicit expression for I_1 is found. This expression is further simplified by assuming an atmosphere with large radius. The resulting expressions for I_1 closely approximate the total intensity of scattered light for small values of λ , the albedo of the scattering particle, or for small values of θ , the angle between the direction of light incident on the planet and the ray directed toward the observer. It is further pointed out that entirely different expressions are found for I_1 when the atmosphere is assumed to consist of plane and parallel layers. Orig. art. has: 43 formulas and 3 figures.

ASSOCIATION: None

SUBMITTED: 20Feb63

DATE ACQ: 26Dec63

ENCL: 00

SUB CODE: AS

NO REF SOV: 001

OTHER: 001

3/3

Card

MININ, I.N.

Diffusion of radiation in a semi-infinite medium with nonisotropic scattering. Part 2. Vest. LGU 18 no.13:106-118 '63. (MIRA 16:9)

(Radiation) (Scattering (Physics))

L 11192-63

Pe-4/Pi-4--GW

EW(1)/FCC(w)/BDS/ES(v)--AFFTC/ASD/ESD-3/APGC/SSD--

ACCESSION NR: AP3001243

8/0033/63/040/003/0496/0503

AUTHOR: Minin, I.N.; Sobolev, V.V.

TITLE: Contribution to the theory of the scattering of light in planetary atmospheres ⁶⁹₆₈

SOURCE: Astronomicheskii zhurnal, v. 40, no. 3, 1963, 496-503

TOPIC TAGS: planetary atmosphere, scattering of light, luminosity of planetary atmosphere, twilight phenomena, terminator

ABSTRACT: This theoretical paper examines the problem of the scattering of light in a spherical atmosphere, continuing and extending the investigation reported in the authors' paper in "Iskusstvennyye sputniki Zemli (Artificial Earth satellites)", no. 14, Izd-vo AN SSSR, Moscow, 1962, in which the problem is approximately reduced to a certain differential equation. In the present paper the problem is reduced to an integral equation. The solution of this problem is essential for the study of the luminosity of a planet in the vicinity of the terminator, i.e., that region of the planet in which the altitude of the sun over the horizon is low, also for the construction of a theory of twilight phenomena. The integral equation for the source function is developed on the premise of

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isotropic scattering of the light. For the sake of simplicity, the planetary atmosphere is imagined to consist of plane-parallel layers. However, it is assumed that these layers, in a given locality, are illuminated by the solar rays as though they were part of a spherical atmosphere. The reflection of the light from the planetary surface is taken into account. If it is assumed that the atmospheric layers are illuminated by parallel solar rays at each point, then the equation obtained thereby yields the well-known equation of the theory of the scattering of radiation in a planetary atmosphere. The integral equations obtained in the present paper will subsequently be numerically solved for various cases. In particular, the authors intend to examine in detail the case of a gaseous atmosphere in which the absorption coefficient decreases exponentially with elevation, also the case of a two-layer atmosphere consisting of a lower cloud-filled layer and an upper gaseous layer. The results of the calculation will be applied to the study of the luminosity of the atmospheres of the Earth and other planets when the sun is at a low local altitude. Here the first-order scattering will be taken into account exactly, the higher-order scattering approximately. It is further intended to generalize the results of this study. There are 46 numbered equations and 2 figures.

ASSOCIATION: Astronomicheskaya observatoriya Leningradskogo gos. universiteta

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EWI(1)/FCC(w)/BDS/ES(v)--- AFFTC/ESD-3 Pa-4 GW

PHASE I BOOK EXPLOITATION

SOV/6434

Gorbatskiy, V. G., and I. N. Minin

Nestatsionarnyye zvezdy (Unstable Stars) Moscow, Fizmatgiz, 1963. 355 p.
(Series: Problemy teoreticheskoy astrofiziki) 2000 copies printed.

Editorial Board of the Series: V. A. Ambartsumyan, E. R. Mustel', A. B. Severnyy, and V. V. Sobolev; Ed.: G. S. Kulikov; Tech. Ed.: I. Sh. Aksel'rod.

PURPOSE: This book is intended for astronomers and astrophysicists.

COVERAGE: Unstable stars, including novae, supernovae, Wolf-Rayet, and Be-types, are investigated on the basis of their emission characteristics. Shell dynamics during flareup are examined. The instability of the stars is interpreted chiefly on the basis of the structure of the outer layers, since little data is available on the interior of such stars. Some attention is given to the application of gasdynamics and electrodynamic techniques

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L 10339-63
Unstable Stars

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to the study of unstable stars. Cepheid variables are not examined. Chs. 1, 3, 7, 9, 10, and para. 19 were written by V. G. Gorbatskiy and Chs. 2, 4, 5, 6, and 7, by I. N. Minin. The authors thank colleagues at the Department of Astrophysics of Leningrad University and the Department of Stellar Physics of the Crimean Astrophysical Observatory. There are 293 references, including 132 Soviet.

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Foreword

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PART I. NOVAE

Ch. 1. The Qualitative Interpretation of Observational Data

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1. Typical novae

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2. Qualitative interpretation of a nova outburst

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3. Characteristics of individual novae

28

Card 2/5 ✓

MININ, I. N.; SOBOLEV, V. V.

"Light scattering in the spherical atmosphere."

paper presented at the Atmospheric Radiation Symp, Leningrad, 5-12 Aug 64.

MININ, I.N.; PILIPOSYAN, A.G.; SHIDLOVSKAYA, N.A.

Tables of Ambartsumian's functions for anisotropic scattering.
Uch. Zap. LGU no.323:12-36 '64. (MIRA 17:12)

ACCESSION NR: AP4043498

S/0293/64/002/004/0610/0618

AUTHOR: Minin, I. N., Sobolev, V. V.

TITLE: Light scattering in a spherical atmosphere. Part III

SOURCE: Kosmicheskiye issledovaniya, v. 2, no. 4, 1964, 610-618

TOPIC TAGS: planetary atmosphere, light scattering, atmospheric optics, atmospheric absorption coefficient, planet brightness, planetary albedo

ABSTRACT: In this article, as in the previous parts of their study (Iskusstvenny*ye sputniki Zemli, No. 14, Izd-vo AN SSSR, 1962, p. 7; Kosmicheskiye issledovaniya, 1, No. 2, 227, 1963), the authors consider the problem of diffusion of radiation in a planetary atmosphere illuminated by the sun's rays. The curvature of atmospheric layers is taken into account. In the earlier studies the principal equations of the problem were derived and a solution was found for a case when the absorption coefficient for the atmosphere is constant. In this third part of the study the assumption is made that the absorption coefficient decreases exponentially with height. The problem is solved in the first approximation and the following computations were made: 1. brightness of the planet near the terminator, and 2. brightness of the zenith during observations from the

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ACCESSION NR: AP4043498

earth's surface for different zenith distances of the sun. Table 2 in the original gives the brightness of a planet near the terminator. Table 3 gives the values I_0 and ΔI (where I_0 is the intensity caused by first-order scattering in the case of a spherical indicatrix of scattering and ΔI is the intensity caused by scattering of higher orders) as a function of solar zenith distance Ψ for different values of the optical thickness τ_0 of the atmosphere. The value ΔI is given for two values of the albedo of a planetary surface ($A = 0.2$ and $A = 0.8$), approximately corresponding to summer and winter conditions. These data show that the relative role of higher-order scattering changes little with a change in solar zenith distance. Table 4 gives the values of the total brightness of the zenith. A comparison of computed and observed values of zenith brightness shows good agreement. The presented theory of light scattering in a spherical atmosphere is rather approximate, but it can be made more precise by taking into account a term neglected in one of the formulas or by using an integral equation describing diffusion of radiation in a spherical atmosphere derived earlier by the authors (Astron. zh., 40, No. 3, 496, 1963). The radiation transport equation used does not take into account the refraction of radiation. However, refraction apparently must be taken into account only in a study of first-order scattering for angles Ψ close to $\pi/2$. In a study of higher-order scattering refraction probably can be

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ACCESSION NR: AP4043498

neglected, as it is neglected in the ordinary theory of light scattering in planetary atmospheres. "The authors wish to thank Ye. B. Babkova and L. P. Savitskaya for computations involved in this study. Orig. art. has 48 formulas and 4 tables.

ASSOCIATION: none

SUBMITTED: 31Jan64

ENCL: 00

SUB CODE: AA, OP

NO REF SOV: 007

OTHER: 001

3/3

Card

MININ, I.N.

Light scattering in dust nebulae. Astron. zhur. 41 no.4:
662-668 J1 -Ag '64 (MIRA 17:8)

1. Leningradskiy gosudarstvennyy universitet im. A.A. Zhdanova.

L 27899-65 EWT(1) IJP(c)

ACCESSION NR: AP4016502

S/0020/64/154/005/1059/1062

AUTHOR: Minin, I.N.

TITLE: On the unsteady luminosity of a semi-infinite medium

SOURCE: AN SSSR. Doklady*, v. 154, no. 5, 1964, 1059-1062

TOPIC TAGS: luminosity, semiinfinite medium, light scattering,²¹ isotropic scattering, quantum, quantum survival, radiation intensity, unsteady diffusion, light reflection, Laplacian transformation, light quantum, optical depth

ABSTRACT: The unsteady luminosity of a homogeneous semi-infinite medium in which the isotropic scattering of light provides for the probable "survival" of the quantum has been studied. A method is proposed for solving various problems pertaining to the theory of unsteady diffusion of radiation which is based on the assumption that the optical properties of the medium do not change in the course of time. The method involves finding the Laplacian transformation in any point of the radiation field from a corresponding

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L 27899-65

ACCESSION NR: AP4016502

point in a steady-state case. In two cases under consideration, one included diffused reflection of light in a semi-infinite medium illuminated by parallel beams, and the other a medium illuminated by evenly distributed sources. The solution to the problem in a steady-state case as made by V.A. Ambartsumyan was as follows:

$$I(\eta) = B_0 \frac{\varphi(\eta)}{\sqrt{1-\eta^2}} \quad (22)$$

where $B_0 dt$ represents the quantity of energy emitted by the sources located in an elementary volume with a cross section of 1 cm^2 and optical depth of dt per 1 sec.; $I(\eta)$ is the intensity of radiation emerging from the medium and comprising an angle $\arccos \eta$ with the normal (line) toward the boundary. Orig. art. has: 32 formulas.

ASSOCIATION: Leningradskiy gosudarstvennyy universitet imeni A.A. Zhdanova (Leningrad State University)

SUBMITTED: 12Oct63

ENCL: 00

SUB CODE: OP, NP

NO REF SOV: 004

OTHER: 001

Card 2/2

MININ, I.N.

Some calculations of light scattering in dust nebulae. Trudy
Astrofiz. inst. AN Kazakh. SSR 5:258-261 '65.

(MIRA 18:6)

L 5431-66 EWT(1)/FCC GW

ACC NR: AT5026206

SOURCE CODE: UR/2703/65/000/328/0039/0043

AUTHOR: Minin, I. N. ^{44,55}

ORG: Astronomical Observatory, Leningrad State University ^{44,55} (Astronomicheskaya observatoriya, Leningradskiy gosudarstvennyy universitet)

TITLE: On the scattering of light ^{12,44,55} in planetary atmospheres

SOURCE: Leningrad. Universitet. Uchenyye zapiski, no. 328, 1965. Seriya matematicheskikh nauk, no. 39. Trudy Astronomicheskoy observatorii, v. 22, 39-43

TOPIC TAGS: light scattering, Rayleigh scattering, planetary atmosphere, aerosol, light reflection, light polarization, astrophysics

ABSTRACT: The scattering of light off a plane layer with finite optical thickness τ_0 is considered. The layer is bounded below by a reflecting floor and is illuminated with parallel rays incident at a given angle to the normal. The scattering of this light is calculated as arising from molecules (Rayleigh scattering) and from atmospheric aerosols. The motivation for this calculation

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L 5431-66

ACC NR: AT5026206

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proceeds from investigation of optical properties of planetary atmospheres, particularly those of earth and Mars. In this work, account is taken of the light polarization in reflection off the atmospheric floor and scattering by the aerosols. Approximate formulas describing the radiation emerging from the planetary atmosphere are obtained. Orig. art. has: 18 formulas.

SUB CODE: AA, ES, OP/ SUBM DATE: 00/

ORIG REF: 003/

OTH REF: 002

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Card 2/2

MININ, I.N.

Light scattering in a one-dimensional nonsteady-state medium.
Astrofizika 1 no.2:173-181 Ja. '65. (MIRA 18:10)

1. Leningradskiy gosudarstvennyy universitet.

Минин, И. С.
MININ, I.S.; MOROV, N.S.

Machine for cold bending 3 to 6-inch pipes. Rats. 1 izobr. predl.
v stroi. no. 72:12-13 '54. (MIRA 10:9)
(Pipe bending)

SHTERN, I.A.; KIPNIS, Yu.B.; PLOTNIKOV, I.V.; PAVLOV, S.A.; PAVLOV, N.N.;
VTOROV, G.N.; PROKURAT, R.E.; GLAGOLEVA, K.I.; KOCHERZHINSKAYA,
Ye.L.; FEDOROVA, L.V.; MININ, I.T.

Artificial carbocylate leather. Kozh.-obuv. prom. 6
no.2:32-34 F'64. (MIRA 17:5)

MININ, K.

School principal is elected by a collective. Sov.profsoiuzy
18 no.22:19-20 N '62. (MIRA 15:12)
(School superintendents and principals)

S/120/62/000/002/011/047
E039/E520

21.6000
AUTHORS: Sidorenko, V.V., Ivanov, V.P. and Minin, K.F.
TITLE: A gamma-dosimeter with a gas multiplication counter
and a pulsed supply system
PERIODICAL: Pribery i tekhnika eksperimenta, no.2, 1962, 55-58
TEXT: This instrument fills the need for a single detector
to cover a wide range of dose rates (0.05 to 1000 r/hr). The
probe unit contains a gas multiplication counter СИ-3БП (SI-3BG)
and blocking generator СП15П (6P15P) in an aluminium cylinder
(65 mm diameter and 260 mm high; weight 620 g). The control unit,
dimensions 180 x 145 x 205 mm³, weighs 3 kg and uses a СБ-1М/100
(SB-1M/100) electromechanical counter. A calibration obtained for ✓
dose rates up to 1200 r/hr with a Co⁶⁰ source showed that the
indicated dose agreed with the calculated value to $\pm 5\%$. The
sensitivity is not less than 0.05 r/hr. For changes of $\pm 10\%$ in
the supply voltage the readings change by not more than $\pm 4\%$.
There is practically no background count-rate. For temperature
changes of $+50$ to -40°C the readings change by not more than $\pm 5\%$.
The probe can be used at distances of up to 50 m from the control

Card 1/2

A gamma-dosimeter with a gas ...

S/120/62/000/002/011/047
E039/E520

unit. A detailed description of the circuit is given. There are
5 figures.

SUBMITTED: August 4, 1961

✓B

Card 2/2

IVANOV, V.P.; MININ, K.F.; KUZIN, A.M.

Wide-range roentgenometer. Prib. i tekhn. eksp. 8 no.5:65-69
S-0 '63. (MIRA 16:12)

24.7500

67187

SOV/58-59-7-15482

Translation from: Referativnyy Zhurnal Fizika, 1959, Nr 7, p 123 (USSR)

AUTHORS: Minin, L.P., Terminasov, Yu.S.

TITLE: X-Ray Diffraction Study of Aluminum Deformed by Stretching at Room Temperature and Low Temperature

PERIODICAL: Uch. zap. Leningr. gos. ped. in-ta im. A.I. Gertsena, 1958, Vol 141, pp 225 - 232

ABSTRACT: The authors studied the substructure of Al that had been deformed by stretching at temperatures ranging from 20° to -194°C. A special attachment made it possible to effect the deformation of the sample and obtain X-ray photographs both at room temperature and at low temperature (-194°C). The X-ray photography was effected by the reverse exposure method. It was established that there exists an interconnection between lattice distortions and the process of crushing of the blocks that takes place under plastic deformation. The increase of stresses in the initial stage of deformation paves the way for the process of crushing. The crushing of the blocks, as well as a certain disorientation that they

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SOV/58-59-7-15482

X-Ray Diffraction Study of Aluminum Deformed by Stretching at Room Temperature and Low Temperature

undergo, in turn promotes a further increase of distortions. It is probably the elastic stresses resulting from cooling that pave the way for the intensive crushing of the blocks that takes place when Al is deformed at low temperatures.

Card 2/2

NIKITIN, S.P., prof.; MININ, L.S., st. prepod., red.

[Laboratory manual on the strength of materials] Laboratornyi praktikum po soprotivleniiu materialov. 3. izd. Moskva, Mosk. energet. inst. 1964. 115 p.
(MIRA 18:12)

YERDAKOV, Vadim Ivanovich, inzh.; MININ, Leonid Sergeyevich, inzh.;
TIKHOMIROV, Ye.N., prof., retsenzents; DARKOV, A.V., doktor
tekhn. nauk, retsenzents; SAPOZHNIKOV, N.M., inzh., nauchnyy
red.; KOPTEVSKIY, D.Ya., red. izd-va; YEZHOVA, L.L., tekhn.
red.

[Laboratory practical work on the strength of materials] La-
boratornyi praktikum po soprotivleniiu materialov dlia studentov
zaochnykh vtuzov. Moskva, Gos. izd-vo "Vysshaya shkola," 1961.
188 p. (MIRA 15:4)
(Strength of materials--Testing) (Testing machines)

MININ, L.S., starshiy prepodavatel'

Dynamic strength of end-type electrical machines with printed
circuits. Trudy MEI no.38:291..304 '62. (MIRA 17:2)

ITSKOVICH, G.M.; VINOKUROV, A.I.; Primal uchastiye:
MININ, I.S.; MAKUSHIN, V.M., laureat Leninskoy premii,
prof., retsenzent; SHPIRO, G.S., kand. tekhn.nauk, nauchn.
red.; BORODINA, N.N., red.; CHIZHEVSKIY, E.M., tekhn.red.

[Manual for solving problems on the strength of materials]
Rukovodstvo k resheniiu zadach po soprotivleniiu materialov.
Moskva, Rosvuzizdat, 1963. 351 p. (MIRA 16:8)
(Strength of materials--Problems, exercises, etc.)

MAKLAKOV, I.A.; MININA, L.S.

Height of the tropopause over Moscow in the period from
1957 to 1960. Trudy TSIP no.137:21-26 '64.

Fluctuations of tropopause altitude in connection with
the temperature variations in the troposphere and lower
stratosphere. Ibid.:37-43 (MIRA 17:9)

MININA, L.S.; BATYAYEVA, T.F.

Structure of the atmosphere following the invasion of tropical
air over Europe. Trudy TSIP no.137:44-53 '64.

(MIRA 17:9)

BATYAYEVA, T.F.; MININA, L.S.

Weather and the circulation of the atmosphere in the winter
of 1962-1963. Trudy TSIP no.137:151-159 '64. (MIRA 17:9)

MININ, M.I.

"On the Road to Raising State Farm Production."

USSR Home Service. 12 April 1955.

MININ, Mikhail Kuz'mich; SORKIN, S., red.; PAVLOVA, S., tekhn.red.

[New wage system at the "Smugorovo" State Farm] Novoe v opiate
truda v sovkhose "Smugorovo." Moskva, Mosk.rabochii, 1960.
47 p. (MIRA 13:5)

1. Direktor sovkhosa "Smugorovo", Moskovskoy oblasti (for Minin).
(State farms) (Wages)

KAGAN, Iosif Zakharovich; MININ, M.N., red.; GARMASH, L.M., otv.za vypusk;
SUKHAREVA, R.A., tekhn.red.

[Introduction of electric slag welding; "Penzkhimmash" Plant of
the Penza Economic Council] Opyt vnedrenia elektroshtakovoi
svarki; zavod "Penzkhimmash" Penzenskogo sovnarkhoza. Moskva,
1958. 16 p. (Moskovskii dom nauchno-tekhnicheskoi propagandy.
Peredovoi opyt proizvodstva. Seriya: Tekhnologiya mashinostro-
eniia, no.30. Svarka, psika i metallizatsiia).

(MIRA 13:10)

(Penza Province--Electric welding)

MININ, M.S. Ansh.

Experience in producing gas turbine blades at the "Economizer"
plant. *Energomashinostroenie* 4 no.12:41,48 D '58.

(MIRA 11:12)

(Gas turbines)

AUTHOR: Minin, M.S., Engineer SOV/117-58-11-16/36

TITLE: The Manufacture of Gas Turbine Blades (Izgotovleniye gazoturbinnykh lopatok)

PERIODICAL: Mashinostroitel', 1958, Nr 11, pp 19 - 20 (USSR)

ABSTRACT: The principal problem in the manufacture of gas turbine blades is the selection of basic surfaces from which the machining is to proceed. Steam turbine and compressor blades cannot be compared with gas turbine blades in this respect. The profile surface of gas turbine blades is very complex, and there are various methods for the selecting basic surfaces (Figure 1 and 2). The mechanical processing of these blades is done in 50-60 operations. The back edge of the blade is machined on a lathe by using a copy. The different operations are described. There are no special machines for the manufacture of gas turbine blades. There are 2 sets of diagrams.

1. Gas turbine blades--Production 2. Machine tools---Performance

Card 1/1

MININ, M.Ye.

Adoption of charging bunkers and their role in speeding up the
mechanization of winning machine peat in White Russia. Sbor.
nauch.trud.Bel.politekh.inst. no.65:83-89 '59.

(MIRA 13:5)

(White Russia--Peat machinery)

NARYSHKIN, I.I.; MININ, N.A.

Polarography of melts over lithium and potassium chlorides using a
lead dropping electrode. Zhur.prikl.khim. 34 no.10:2353-2356 0
'61. (MIRA 14:11)

(Salts) (Polarography) (Electrodes, Lead)

MININ, N. D.

Drifting by cutter-loader in the Moscow Basin. Mast. ugl. 3
no. 12:7-8 D '54. (MLRA 8:6)

1. Brigadir prokhodchikov shakhty no. 67 "Zhdankovskaya"
kombinata Tulaugol'.
(Moscow Basin--Coal mines and mining) (Coal mining
machinery)

MININ, N. D.

We shall not stop with the achievements made. Mast. ugl. 4 no. 8:14
Ag '55. (MLRA 8:10)

1. Mashinist prokhodcheskogo kombayna shakhty no. 67 kombinata
Tulaugol'

(Tula Province--Coal miners)

MININ, Nikolay Dmitriyevich; POPEKOV, Boris Ivanovich; KOLOMIYTSSEV, A.D.,
~~otvetstvennyy~~ redaktor; MADNINSKAYA, A.A., tekhnicheskiy redaktor

[Gamma ray relays for the automatization in the coal industry]
Gamma-rele dlia avtomatizatsii v ugol'noi promyshlennosti. Moskva,
Ugletekhnizdat, 1956. 63 p. (MLRA 9:7)
(Gamma rays--Industrial application)
(Coal mining machinery)

MININ, N.F.

At the Moscow Exhibition of Fruit Culture. Zashch. rast. ot vred.
1 bol. 8 no.2:57 F '63. (MIRA 16:7)

1. Chlen byuro seksii sadovodstva Vserossiyskogo obshchestva okhrany
prirody.

(Moscow—Agricultural exhibitions)

MININ, N. I.

USSR / Pharmacology, Toxicology, Analgesics.

V

Abs Jour : Ref Zhur - Biol., No 20, 1958, No 94189

Authors : Minin, N. I.; Starchenko, N. N.

Inst : Moscow Medical Institute

Title : Scophedal (Scopolamine - Eikodal - Ephetonin)
as Basic Narcosis in the Surgical Clinic.

Orig Pub : Tr. 1-go Mosk. med. in-ta, 1957, 3, 119-124.

Abstract : Combined hypodermic anaesthetization (1 ml of scophedal (I), three minutes before the operation under local anaesthesia) was applied to 50 patients during different operations. From 10 - 20 minutes after the injection of I in the majority of patients this resulted in a state of general quiescence, a feeling of fatigue, indifferent attitude towards the surroundings, dizziness, dryness in the mouth, and a drowsy

Card 1/2

MININ, N.I.; STARCHENKO, N.N.

Intravenous use of scopolamine-morphine-caffeine mixture as the principal anesthetic. Trudy 1-go MMI 3:125-128 '57. (MIRA 14:5)

(ANESTHETICS)

MININ, N.I., dotsent; BABIN, V.B.; KOFMAN, I.L.; MANEVICH, V.A.;
MIKHEL'SON, V.A.; YUREVICH, V.M.

Concentration of ether in the blood during various types of
ether-oxygen anesthesia. Vest.khir. 85 no.9:95-100 8 '60.

(MIRA 13:11)

1. Iz fakul'tetskoy khirurgicheskoy kliniki (zav. - prof. I.S.
Zhorov) sanitarno-gigiyenicheskogo fakul'teta 1-go Moskovskogo
ordena Lenina meditsinskogo instituta imeni I.M. Sechenova.
(ETHER (ANESTHETIC))

MARUSENKO, Yakov, Il'ich; ZEMTSOV, Aleksey Anisimovich; SEMLYANSKAYA, Lidiya Pavlovna; PANKOV, Arkadiy Mikhaylovich; MININ, Nikolay Kondrat'yevich; MORDOVINA, L.G., tekhn. red.

[Hydrography of Western Siberia] Gidrografiia Zapadnoi Sibiri. Tomsk, Izd-vo Tomskogo univ. Vol.1. [General characteristics of waters] Obshchaia kharakteristika vod. 1961. 169 p.

(MIRA 14:11)

(Siberia, Western—Hydrography)

MININ, N. P.

24288

MININ, N. P. K voprosu o vzaimosvyazyakh i glubokikh limfaticheskikh sosudov nizhney knochnosti. Trudy Leningr. San.-Gigien. med. in-ta, T. III, 1949, S. 134-49. - Bibliogr: 19 nazv.

SO: Letopis, No. 32, 1949.

MININ, N.P.

MININ, N.P., assistant

Anatomy of superficial lymph vessels of the lower extremity.
Trudy ISGMI 9:158-163 '51. (MIRA 11:1)

1. Kafedra normal'noy anatomii Tomskogo meditsinskogo instituta
im. V.M.Molotova (nacunyy rukovoditel' - chl-korr. AMN SSSR prof.
Zhdanov D.A.)
(LYMPHATICS) (EXTREMITIES, LOWER)

MININA, Ol'ga Mikhaylovna; KRUG, Ye.K., kand. tekhn. nauk,
otv. red.; LETNEV, B.Ya., red. izd-va; LAUT, V.G., tekhn. red.

[Determination of the dynamic characteristics and parameters
of standard control objects] Opređenje dinamicheskikh kha-
rakteristik i parametrov tipovykh reguliruemyykh ob"ektov.
Moskva, Izd-vo AN SSSR, 1963. 44 p. (MIRA 16:10)
(Automatic control)

MININ, P. I.

Research in the drawing of steel bars. Moskva, Gos. nauch.-tekhn. izd-vo
mashinostroit. lit-ry, 1948. 81 p. (49-29335)

TS320.M635

MININ, P. I., inzh. (Kazan')

Sand and gravel quarries located in rivers. Put' 1 put. khoz. no. 3:33-74
Mr '58. (MIRA 11:4)

(Tatar A.S.S.R.--Quarries and quarrying)

MININ, P.I., inzh.

Experience in riveting thick components in spans. Transp.
stroil. 8 no.8:30-31 Ag '58. (MIRA 11:10)
(Bridges, Iron and steel) (Rivets and riveting)

MININ, P.I., insh.

Device for checking dimensions of tunnels. Transp. stroi. 9 no.11:
42-44 N '59 (MIRA 13:3)
(Gauges) (Tunneling)

MININ, P.I., inzh. (Kazan')

Regulation of water flow under small bridges. Put' i put.khoz.
no.12:9 D '59. (MIRA 13:4)
(Rivers--Regulation) (Railroad bridges)

MININ, P.I., insh. (Kazan')

Practices in regulating the passage of spring floods under
bridges. Put' 1 put.khoz. 4 no.3:11 Mr '60.

(MIRA 13:5)

(Flood control)

MININ, P. I., CAND TECH SCI, ^{II} METHODS OF INSTRUMENTAL
OBSERVATIONS OF THE TECHNICAL CONDITION OF TUNNELS AND
LARGE-SCALE MINING. ^{//} MOSCOW, 1961. (MIN OF HIGHER AND
SEC SPEC ED RSFSR. MOSCOW MINING INST IM I. V. STALIN).
(KL, 2-61, 210).

-158-

MININ, P.I., inzh.

Apparatus for surveying the inside of tunnels. Gor. zhur. no.4:
68-69 Ap '61. (MIRA 14:4)

1. Kazanskaya laboratoriya Vsesoyuznogo nauchno-issledovatel'skogo
instituta transportnogo stroitel'stva.
(Mine surveying--Equipment and supplies)

MININ, P.I., kand.tekhn.nauk; GRITSYK, V.I., inzh.; KHRAMOV, M.G., inzh.

Stabilizing the banks of a dirt roadbed by planting grass. Transp.
stroil. ll no.4:34-36 Ap '61. (MIRA 14:5)
(Kazan--Railroads--Earthwork) (Soil binding)

MININ, P.I., kand.tekhn.nauk

Polar method of laying out the borders of the edges of cuttings
on steel hillsides. Transp. stroi. 12 no.8:41-42 Ag '62.
(Geodesy) (Earthwork)

MININ, P.I., kand.tekhn.nauk

Instrument observations for the opening of the planned profile of a mine. Shakht. stroi. 7 no.1:23-25 Ja '63. (MIRA 16:2)

1. Vsesoyuznyy nauchno-issledovatel'skiy institut okhrany truda
Vsesoyuznogo tsentral'nogo soveta professional'nykh soyuzov, Kazan'.
(Mine surveying)

L 07911-67 EWT(d)/EWT(1) GW
ACC NR: AP6032357 SOURCE CODE: UR/0270/66/000/007/0040/0040

AUTHOR: Minin, P. I.

TITLE: Effect of the shape of bright spots on the precision of determining distances
in range-finder measurements *gm*

SOURCE: Ref. zh. Geodeziya, Abs. 7.52.281

REF SOURCE: Nauchn. raboty in-tov okhrany VTsSPS, vyp. 3(35), 1965, 102-105

TOPIC TAGS: optic range finder, measurement error, distance measurement,
range finder, bright spot

ABSTRACT: An optical diagram of the INIM range finder has been analyzed for
measuring the distance to points of the internal contour line of structures or mine
cavities. Measurements may be carried out in total darkness. The method of light
ranging, i. e., projection of bright spots on the observed point of a contour line, is
used in the range finder. High measurement precision is achieved by means of
rectangular bright spots having sides of 1:5 with the elongated sides positioned in
the line of the transverse profile. The maximum error is 4.95 mm in measuring a
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